

Análisis geométrico de la distancia lorentziana en subvariedades marginalmente atrapadas

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- Obviously, $p \ll q$ implies $p < q$. As usual, $p \leq q$ means that either $p < q$ or $p = q$.
- For a subset $S \subset M$, one defines the **chronological future of S** as

$$I^+(S) = \{q \in M : p \ll q \text{ for some } p \in S\},$$

and the **causal future of S** as

$$J^+(S) = \{q \in M : p \leq q \text{ for some } p \in S\}.$$

Thus $S \cup I^+(S) \subset J^+(S)$.

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- As a matter of fact, **globally hyperbolic** spacetimes¹ turn out to be the natural class of spacetimes for which the Lorentzian distance function is **finite-valued and continuous**.

¹ M is globally hyperbolic provided (a) $J^+(p) \cap J^-(q)$ is compact for any $p, q \in M$, and (b) M is causal, i.e., there are no closed causal curves in M .

Lorentzian distance function from a point

- Given a point $p \in M$, one can define the **Lorentzian distance function from p** by $d_p : M \rightarrow [0, +\infty]$

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- Let $T_{-1}M|_p$ be the fiber of the **unit future observer bundle** of M at p , that is,

$$T_{-1}M|_p = \{v \in T_pM : v \text{ is a future-directed timelike unit vector}\}.$$

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- Define the function $s_p : T_{-1}M|_p \rightarrow [0, +\infty]$ by

$$s_p(v) = \sup\{t \geq 0 : d_p(\gamma_v(t)) = t\},$$

where $\gamma_v : [0, a) \rightarrow M$ is the future inextendible geodesic starting at p with initial velocity v .

Lorentzian distance function from a point

- Then, one can define the subset $\tilde{\mathcal{I}}^+(p) \subset T_p M$ given by

$$\tilde{\mathcal{I}}^+(p) = \{tv : \text{for all } v \in T_{-1}M|_p \text{ and } 0 < t < s_p(v)\}$$

and consider the subset $\mathcal{I}^+(p) \subset M$ given by

$$\mathcal{I}^+(p) = \exp_p(\text{int}(\tilde{\mathcal{I}}^+(p))) \subset I^+(p).$$

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Lemma 1 (Erkekoglu, García-Río and Kupeli, 2003)

Let M be a spacetime and $p \in M$.

- 1 If M is **strongly causal** at p^a , then $s_p(v) > 0$ for all $v \in T_{-1}M|_p$ and $\mathcal{I}^+(p) \neq \emptyset$.
- 2 If $\mathcal{I}^+(p) \neq \emptyset$, then the Lorentzian distance function d_p is **smooth** on $\mathcal{I}^+(p)$ and its gradient $\bar{\nabla} d_p$ is a **past-directed timelike (geodesic) unit** vector field on $\mathcal{I}^+(p)$.

^aGiven any neighborhood U of p there is a neighborhood $V \subset U$ of p such that every causal curve segment with endpoints in V is entirely contained in U .

Hessian comparison results for the Lorentzian distance

- For every $c \in \mathbb{R}$, let us define

$$h_c(t) = \begin{cases} \frac{1}{\sqrt{c}} \sinh(\sqrt{c} t) & \text{if } c > 0 \text{ and } t > 0 \\ t & \text{if } c = 0 \text{ and } t > 0 \\ \frac{1}{\sqrt{-c}} \sin(\sqrt{-c} t) & \text{if } c < 0 \text{ and } 0 < t < \pi/\sqrt{-c}. \end{cases}$$

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- Observe that the index of a Jacobi field along a timelike geodesic in a Lorentzian space form of constant curvature c is given by

$$I_{\gamma_c}(J_c, J_c) = -\frac{h'_c(t)}{h_c(t)} \langle x, x \rangle.$$

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$$I_{\gamma_c}(J_c, J_c) = -\frac{h'_c(t)}{h_c(t)} \langle x, x \rangle.$$

- On the other hand, $\frac{h'_c(t)}{h_c(t)}$ is the future mean curvature of the level set

$$\Sigma_c(t) = \{q \in \mathcal{I}^+(p) : d_p(q) = t\} \subset M_c^n.$$

Hessian comparison results for the Lorentzian distance

Lemma 2 (Alías, Hurtado, Palmer, 2010)

Let M be a spacetime such that $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M . Assume that there exists a point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and let $q \in \mathcal{I}^+(p)$ (with $d_p(q) < \pi/\sqrt{-c}$ when $c < 0$). Then for every spacelike vector $x \in T_q M$ orthogonal to $\overline{\nabla} d_p(q)$

$$\overline{\nabla}^2 d_p(x, x) \leq -\frac{h'_c}{h_c}(d_p(q))\langle x, x \rangle,$$

where $\overline{\nabla}^2$ stands for the Hessian operator on M .

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- The proof of Lemma 2 follows from the fact that

$$\bar{\nabla}^2 d_p(x, x) = I_\gamma(J, J)$$

where γ is the radial future directed unit timelike geodesic from p to q and J is the Jacobi field along γ with $J(0) = 0$ and $J(s) = x$, and is strongly based on the maximality of the index of Jacobi fields.

Hessian comparison results for the Lorentzian distance

- On the other hand, under the assumption that the sectional curvatures of the timelike planes of M are bounded **from above** by a constant c , we get the following result.

Lemma 3 (Alías, Hurtado, Palmer, 2010)

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- The proof is similar to that of Lemma 2.

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- Our first objective is to compute the Hessian of u . To do that, observe that

$$\bar{\nabla} r = \nabla u + (\bar{\nabla} r)^\perp$$

along Σ , where $\nabla u = (\bar{\nabla} r)^\top$ stands for the gradient of u on Σ and $(\bar{\nabla} r)^\perp$ denotes the normal component of $\bar{\nabla} r$.

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- By Gauss and Weingarten formulae we get

$$\bar{\nabla}_X \bar{\nabla} r = \nabla_X \nabla u - A_{(\bar{\nabla} r)^\perp} X + \text{II}(X, \nabla u) + \nabla_X^\perp (\bar{\nabla} r)^\perp,$$

for every tangent vector $X \in T\Sigma$, where II denotes the second fundamental form of the submanifold and, for every normal vector η , A_η denotes the Weingarten endomorphism with respect to η .

- It follows from here that

$$\nabla^2 u(X, Y) = \overline{\nabla}^2 r(X, Y) + \langle \text{II}(X, Y), \overline{\nabla} r \rangle$$

for every tangent vector fields $X, Y \in T\Sigma$, where $\overline{\nabla}^2 r$ and $\nabla^2 u$ stand for the Hessian of r and u in M and Σ , respectively.

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- Tracing this expression, one gets that the Laplacian of u is given by

$$\Delta u = \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + m \langle \mathbf{H}, \bar{\nabla} r \rangle,$$

where $\{E_1, \dots, E_m\}$ is a local orthonormal frame on Σ , and

$$\mathbf{H} := \frac{1}{m} \text{tr}(\text{II}) = \frac{1}{m} \sum_{i=1}^m \text{II}(E_i, E_i)$$

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- Consider the function $v = \phi_c(u)$, where $\phi_c(t)$ is a primitive of $h_c(t)$:

$$\phi_c(t) = \begin{cases} \frac{1}{c} \cosh(\sqrt{c} t) & \text{if } c > 0 \text{ and } t > 0 \\ \frac{t^2}{2} & \text{if } c = 0 \text{ and } t > 0 \\ \frac{1}{c} \cos(\sqrt{-c} t) & \text{if } c < 0 \text{ and } 0 < t < \pi/\sqrt{-c}. \end{cases}$$

- Then, the Laplacian of v is given by

$$\begin{aligned}\Delta v &= \phi'_c(u)\Delta u + \phi''_c(u)|\nabla u|^2 \\ &= h_c(u) \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + m h_c(u) \langle \mathbf{H}, \bar{\nabla} r \rangle + h'_c(u)|\nabla u|^2.\end{aligned}$$

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- Assume now that $K_M(\Pi) \geq c$ (resp. $K_M(\Pi) \leq c$) for all timelike planes in M . Then by the Hessian comparison results for r given in Lemma 2 (resp. Lemma 3), one gets that

$$\bar{\nabla}^2 r(X, X) \leq (\geq) - \frac{h'_c}{h_c}(u)(1 + \langle X, \nabla u \rangle^2)$$

for every unit tangent vector field $X \in T\Sigma$.

- Then, the Laplacian of v is given by

$$\begin{aligned}\Delta v &= \phi'_c(u)\Delta u + \phi''_c(u)|\nabla u|^2 \\ &= h_c(u) \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) + mh_c(u)\langle \mathbf{H}, \bar{\nabla} r \rangle + h'_c(u)|\nabla u|^2.\end{aligned}$$

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for every unit tangent vector field $X \in T\Sigma$.

- Therefore,

$$h_c(u) \sum_{i=1}^m \bar{\nabla}^2 r(E_i, E_i) \leq (\geq) - h'_c(u)(m + |\nabla u|^2),$$

which, jointly with the expression above, gives the following inequality for the Laplacian of v

$$\Delta v \leq (\geq) - mh'_c(u) + mh_c(u)\langle \mathbf{H}, \bar{\nabla} r \rangle.$$

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where $v = \phi_c(u)$ and u is the Lorentzian distance function of M restricted on the spacelike submanifold Σ .

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- $K_M(\Pi) \leq c$ implies that

$$\Delta v \geq -mh'_c(u) + mh_c(u)\langle \mathbf{H}, \bar{\nabla} r \rangle.$$

where $v = \phi_c(u)$ and u is the Lorentzian distance function of M restricted on the spacelike submanifold Σ .

- Before stating our main results, we need to introduce some terminology about:
 - Our analytical tool: the weak maximum principle.
 - Our geometric objects: trapped submanifolds.

The Omori-Yau maximum principle

- Following the terminology introduced by Pigola, Rigoli and Setti (2005), the **Omori-Yau maximum principle** is said to hold on an n -dimensional Riemannian manifold Σ if, for any smooth function $u \in \mathcal{C}^2(\Sigma)$ with $u^* = \sup_{\Sigma} u < +\infty$ there exists a sequence of points $\{p_k\}_{k \in \mathbb{N}}$ in Σ with the properties

$$(i) \ u(p_k) > u^* - \frac{1}{k}, \quad (ii) \ |\nabla u(p_k)| < \frac{1}{k}, \quad \text{and} \quad (iii) \ \Delta u(p_k) < \frac{1}{k}.$$

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- Equivalently, for any $u \in \mathcal{C}^2(\Sigma)$ with $u_* = \inf_{\Sigma} u > -\infty$ there exists a sequence of points $\{p_k\}_{k \in \mathbb{N}}$ in Σ satisfying

$$(i) \ u(p_k) < u_* + \frac{1}{k}, \quad (ii) \ |\nabla u(p_k)| < \frac{1}{k}, \quad \text{and} \quad (iii) \ \Delta u(p_k) > -\frac{1}{k}.$$

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- In this sense, the **classical** maximum principle given by Omori (1967) and Yau (1975) states that the Omori-Yau maximum principle holds on every complete Riemannian manifold with **Ricci curvature bounded from below**.

The weak maximum principle

- The **weak maximum principle** is said to hold on Σ if, for any $u \in \mathcal{C}^2(\Sigma)$ with $u^* < +\infty$ there is a sequence $\{p_k\}_{k \in \mathbb{N}}$ in Σ with

$$(i) \quad u(p_k) > u^* - \frac{1}{k}, \quad \text{and} \quad (iii) \quad \Delta u(p_k) < \frac{1}{k}.$$

- Pigola, Rigoli and Setti (2003) proved that the weak maximum principle holds on Σ if and only if Σ is **stochastically complete**.
- Σ is said to be **stochastically complete** if its Brownian motion is stochastically complete, i.e, the probability of a particle to be found in the state space is constantly equal to 1. In other words,

$$\int_{\Sigma} p(x, y, t) dy = 1 \text{ for any } (x, t) \in \Sigma \times (0, +\infty),$$

where $p(x, y, t)$ is the heat kernel of the Laplacian operator².

²For any open $\Omega \subset \Sigma$, $\int_{\Omega} p(x, y, t) dy$ is the probability that a random path starting at x lies in Ω at finite time t . Hence $\int_{\Sigma} p(x, y, t) dy < 1$ means that there is a positive probability that a random path will reach infinity in finite time t .

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- This is equivalent (among other conditions) to the fact that for every $\lambda > 0$, the only non-negative bounded smooth solution u of $\Delta u \geq \lambda u$ on Σ is the constant $u = 0$.
- In particular, every **parabolic** manifold is stochastically complete.

Hence, the **weak max principle** holds on every parabolic manifold.

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Trapped submanifolds in a spacetime

- Following the standard terminology in General Relativity, a spacelike submanifold Σ^m (of arbitrary codimension) of a spacetime M^n is said to be a **future trapped** submanifold if its mean curvature vector field \mathbf{H} is **timelike and future-pointing** everywhere on Σ .

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- Similarly, Σ is said to be **marginally past trapped** if \mathbf{H} is **lightlike and past-pointing** on Σ .
- Finally, Σ is said to be **weakly future trapped** if \mathbf{H} is **causal** (that is, timelike or lightlike) and **future-pointing** everywhere.

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- Finally, Σ is said to be **weakly future trapped** if \mathbf{H} is **causal** (that is, timelike or lightlike) and **future-pointing** everywhere.
- Analogously, Σ is said to be **weakly past trapped** if \mathbf{H} is **causal and past-pointing** on Σ .

Weakly trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \geq c$

Weakly trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \geq c$

Theorem 1

Let M be a spacetime with a reference point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and assume $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M .

- 1 If $c \geq 0$ there exists no stochastically complete, weakly past trapped submanifold contained in $\mathcal{I}^+(p)$.
- 2 If $c < 0$ and Σ is a stochastically complete, weakly past trapped submanifold contained in $\mathcal{I}^+(p) \cap B^+(p, \pi/\sqrt{-c})$, then

$$u_* = \inf_{\Sigma} u \geq \frac{\pi}{2\sqrt{-c}},$$

where u denotes the Lorentzian distance d_p along the hypersurface. In other words, Σ is contained in $B^+(p, \pi/\sqrt{-c}) \cap O^+(p, \pi/2\sqrt{-c})$.

Weakly trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \geq c$

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- Here, for $\delta > 0$, the subsets $B^+(p, \delta)$ and $O^+(p, \delta)$ denote the **future inner ball** and the **future outer ball** of radius δ , that is,
$$B^+(p, \delta) = \{q \in I^+(p) : d_p(q) < \delta\}$$
$$O^+(p, \delta) = \{q \in I^+(p) : d_p(q) > \delta\}.$$

Proof of Theorem 1

- As $K_M(\Pi) \geq c$, we know that

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$$-\frac{1}{k} < \Delta v(p_k) \leq -mh'(u(p_k)) + mh(u(p_k))\langle \mathbf{H}, \overline{\nabla} r \rangle(p_k),$$

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- Observe that, since Σ is weakly past trapped, then

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and, making $k \rightarrow \infty$ here we get $h'_c(u_*) \leq 0$.

- The result then follows by observing that, when $c \geq 0$ then $h'_c(t) > 0$, and if $c < 0$ then $h'_c(t) \leq 0$ when $\pi/2\sqrt{-c} \leq t < \pi/\sqrt{-c}$.

Marginally trapped submanifolds in the chronological future of a point. Case $K_M(\Pi) \geq c$

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Theorem 2

Let M be a spacetime with a reference point $p \in M$ such that $\mathcal{I}^+(p) \neq \emptyset$, and assume $K_M(\Pi) \geq c$, $c \in \mathbb{R}$, for all timelike planes in M . Let Σ be a stochastically complete, marginally trapped submanifold contained in $\mathcal{I}^+(p)$ (with $u_* < \pi/2\sqrt{-c}$ in the case $c < 0$). Then

$$\sup_{\Sigma} |\mathbf{H}_0| \geq \frac{h'_c}{h_c}(u_*),$$

where \mathbf{H}_0 stands for the spacelike component of the lightlike vector field \mathbf{H} which is orthogonal to $\bar{\nabla}r$, and $u_* = \inf_{\Sigma} u$. In particular, if $u_* = 0$ then $\sup_M |\mathbf{H}_0| = +\infty$.

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Corollary 1

Under the assumptions of Theorem 2, if $|\mathbf{H}_0|$ is bounded from above on Σ , then there exists some $\delta > 0$ such that $\Sigma \subset O^+(p, \delta)$, where $O^+(p, \delta)$ denotes the **future outer ball** of radius δ .

Proof of Theorem 2

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- Applying again the weak maximum principle on Σ to the function $v = \phi_c(u)$, with $v_* = \inf_{\Sigma} v = \phi_c(u_*)$, we have

$$-\frac{1}{k} < \Delta v(p_k) \leq -mh'_c(u(p_k)) + mh_c(u(p_k)) \sup_{\Sigma} |\mathbf{H}_0|,$$

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- The last assertion follows from the fact that $h_c(0) = 0$ and $h'_c(0) = 1$.

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Theorem 3

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where \mathbf{H}_0 stands for the spacelike component of the lightlike vector field \mathbf{H} which is orthogonal to $\overline{\nabla} r$, and $u^* = \sup_{\Sigma} u$.

Proof of Theorem 3

- Since $K_M(\Pi) \leq c$ and $\langle \mathbf{H}, \overline{\nabla} r \rangle = |\mathbf{H}_0| > 0$ on Σ , we have

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$$\Delta v \geq -mh'_c(u) + mh_c(u)|\mathbf{H}_0| \geq -mh'_c(u) + mh_c(u) \inf_{\Sigma} |\mathbf{H}_0|.$$

- Applying the weak maximum principle on Σ to the function $v = \phi_c(u)$, with $v^* = \sup_{\Sigma} v = \phi_c(u^*)$, we have

$$\frac{1}{k} > \Delta v(p_k) \geq -mh'_c(u(p_k)) + mh_c(u(p_k)) \inf_{\Sigma} |\mathbf{H}_0|,$$

for $\{p_k\} \subset \Sigma$ with $\lim_{k \rightarrow \infty} v(p_k) = v^*$ and $\lim_{k \rightarrow \infty} u(p_k) = u^*$.

Proof of Theorem 3

- Since $K_M(\Pi) \leq c$ and $\langle \mathbf{H}, \bar{\nabla} r \rangle = |\mathbf{H}_0| > 0$ on Σ , we have

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- Making $k \rightarrow +\infty$ we conclude that

$$\inf_{\Sigma} |\mathbf{H}_0| \leq \frac{h'_c(u^*)}{h_c(u^*)}.$$

Marginally future trapped submanifolds in Lorentzian space forms

- In particular, when the ambient spacetime is a Lorentzian space form, by putting together Theorems 1, 2 and 3 we obtain the following consequence.

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Let M_c^n be a Lorentzian space form of **constant sectional curvature c** and let $p \in M_c^n$. Let Σ be a stochastically complete, marginally trapped submanifold of M_c^n which is contained in $\mathcal{I}^+(p) \cap B^+(p, \delta)$ for some $\delta > 0$ (with $\delta \leq \pi/2\sqrt{-c}$ if $c < 0$). Then

$$\inf_{\Sigma} |\mathbf{H}_0| \leq \frac{h'_c(u^*)}{h_c(u^*)} \leq \frac{h'_c(u_*)}{h_c(u_*)} \leq \sup_{\Sigma} |\mathbf{H}_0|,$$

where $u_* = \inf_{\Sigma} u$ and $u^* = \sup_{\Sigma} u$.

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where $u_* = \inf_{\Sigma} u$ and $u^* = \sup_{\Sigma} u$.

- The estimates are sharp as proved by considering Σ as a constant mean curvature hypersurface of a level set of the Lorentzian distance in M_c^n .

Lorentzian distance function from an achronal hypersurface

- Given $S \subset M^n$ an achronal spacelike hypersurface, one can define the **Lorentzian distance function from S** , $d_S : M \rightarrow [0, +\infty]$, by

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- Then, we can define

$$\tilde{\mathcal{I}}^+(S) = \{t\eta_p : \text{for all } p \in S \text{ and } 0 < t < s(p)\}$$

and consider the subset $\mathcal{I}^+(S) \subset M$ given by

$$\mathcal{I}^+(S) = \exp_S(\text{int}(\tilde{\mathcal{I}}^+(S))) \subset I^+(S),$$

where \exp_S denotes the exponential map with respect to the hypersurface S .

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Lemma 4 (Erkekoglu, García-Río and Kupeli, 2003)

Let S be an achronal spacelike hypersurface in a spacetime M .

- 1 If S is compact and (M, g) is **globally hyperbolic**, then $s(p) > 0$ for all $p \in S$ and $\mathcal{I}^+(S) \neq \emptyset$.
- 2 If $\mathcal{I}^+(S) \neq \emptyset$, then d_S is **smooth** on $\mathcal{I}^+(S)$ and its gradient $\bar{\nabla} d_S$ is a **past-directed timelike (geodesic) unit** vector field on $\mathcal{I}^+(S)$.

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That's all !!
Muchas gracias.